

SOLUTIONS

Module - 1 / JEE-2022

Chemistry	Chapter - 2	Atomic Structure
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EXERCISE-A

$$1. \quad \text{No. of photons} = \frac{E_{\text{Total}}}{E_{\text{per photon}}} = \frac{10^{-17}}{\frac{hc}{\lambda}} = \frac{10^{-17}}{\frac{6.626 \times 10^{-34} \times 3 \times 10^8}{550 \times 10^{-9}}} = 27.67$$

$\Rightarrow$ Min. number of photons = 28

$$2. \quad E_{\text{Photon absorbed}} = E_{\text{Photon 1 emitted}} + E_{\text{Photon 2 emitted}}$$

$$\Rightarrow \frac{hc}{\lambda_{\text{absorbed}}} = \frac{hc}{\lambda_1} + \frac{hc}{\lambda_2} \Rightarrow \frac{1}{300} = \frac{1}{400} + \frac{1}{\lambda_2} \Rightarrow \lambda_2 = 1200 \text{ nm}$$

$$\Rightarrow E_{\text{Photon(2)}} = \frac{1240}{1200} \text{ eV} = 1.03 \text{ eV}$$

$$3. \quad \text{Decelerating potential} = 0.24 \text{ V} \Rightarrow K.E._{\text{Emitted } e^-} \text{ (Photoelectron)} = 0.24 \text{ eV.}$$

Using Einstein's Photoelectric equation: $E_{\text{absorbed}} = K.E._e + W_0 \dots (i)$

$$\Rightarrow E_{\text{absorbed}} = \frac{1240}{300} \text{ eV} = 4.14 \text{ eV}$$

Using (i): $W_0 = 4.14 - 0.24 = 3.9 \text{ eV}$

$$4. \quad W_0 = 7.52 \times 10^{-19} \text{ J}; E_{\text{Photon}} = \frac{hc}{\lambda} = \frac{6.626 \times 10^{-34} \times 3 \times 10^8}{360 \times 10^{-10}} = 5.52 \times 10^{-18} \text{ J}$$

$$\Rightarrow K.E._{\text{Photoelectron}} = (5.52 \times 10^{-18} - 7.52 \times 10^{-19}) \text{ J} = 4.77 \times 10^{-18} \text{ J}$$

Note : $K.E._e$ as calculated using Einstein's Photoelectric eq. is max. possible value. When e^- s are released (knocked out) in photoelectric effect, not all the e^- s have $K.E._{\text{max}}$.

$$\Rightarrow 0 \leq K.E._{\text{any } e^- \text{ from Photoelectric exp.}} \leq K.E._{\text{max}}$$

$$5. \quad \text{No. of } e^- \text{ s} = \frac{\text{Total Charge}}{\text{Charge per } e^-} = \frac{6.39 \times 10^{-19}}{1.6 \times 10^{-19}} = 4$$

6.(B) Cathode Rays – consists of e^- s: e^- s have a fixed e/m ratio.

7.(C) $20W = 20 \text{ J/s}$

$$\begin{aligned} \Rightarrow \text{No. of Photons/sec.} &= \frac{E_{\text{Total/s}}}{E_{\text{Per Photon}}} = \frac{20}{\frac{hc}{\lambda}} = \frac{20}{\frac{6.626 \times 10^{-34} \times 3 \times 10^8}{600 \times 10^{-9}}} = 6.02 \times 10^{19} \text{ per sec} \\ &= \frac{6.02 \times 10^{19}}{6.02 \times 10^{23}} N_{\text{av.}} = 10^{-4} N_{\text{av.}} \end{aligned}$$

8.(D) In Rutherford experiment, α – particles were used.
(He^{2+} ion)

9.(B) $\lambda_{\text{X Rays}} < \lambda_{\text{UV}} < \lambda_{\text{IR}} < \lambda_{\text{Radio}}$

10.(B)
$$\left. \begin{array}{l} \text{I : } h\nu_1 = W_0 + K.E_1 \quad \dots(i) \\ \text{II : } h\nu_2 = W_0 + (K.E_1) \times 2 \quad (\text{given}) \dots(ii) \end{array} \right\} \frac{h\nu_1 - W_0}{h\nu_2 - W_0} = \frac{1}{2} \quad [W_0 = h\nu_0]$$

$\Rightarrow \nu_0 = 8 \times 10^{15} \text{ Hz} = \text{Threshold freq.}$

11.(C) The energy of each photon is :

$$E = \frac{hc}{\lambda} = 3.14 \times 10^{-19} \text{ J.}$$

energy emitted per second by the laser is $5 \times 10^{-3} \text{ J.}$

$$\Rightarrow \text{no. of photons} = \frac{5 \times 10^{-3} \text{ J}}{3.14 \times 10^{-19}} = 1.6 \times 10^{16}$$

12.(C) As we know

$$h\nu = W_0 + K.E.$$

And $K.E \propto \text{stopping potential.}$

$$\text{If } \nu' = 2\nu$$

$\Rightarrow$ stopping potential will become more than double.

EXERCISE-B

1. Here, longest wavelength corresponds to the energy that will just ionize the Li^{2+} ion i.e. ejected e^- has a K.E. = 0

$$\Rightarrow \text{Transition } 2 \longrightarrow \infty \text{ in } \text{Li}^{2+} \text{ ion.} \Rightarrow \frac{1}{\lambda} = 109677 \times 3^2 \times \left(\frac{1}{2^2} - \frac{1}{\infty^2} \right) \Rightarrow \lambda = 405.2 \text{ \AA}$$

2. As explained in the illustrations, Balmer series transition $4 \longrightarrow 2$ in He^+ ion.

$$\Rightarrow \frac{1}{\lambda_{\text{Photon}}} = 109677 \times 2^2 \times \left(\frac{1}{2^2} - \frac{1}{4^2} \right) \quad \dots(i)$$

Let $n_2 \longrightarrow n_1$ be the transition in H-atom.

$$\Rightarrow \frac{1}{\lambda_{\text{Photon}}} = 109677 \times 1^2 \times \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \quad \dots(ii)$$

Comparing (i) and (ii), $n_1 = 1, n_2 = 2 \Rightarrow 2 \longrightarrow 1$ transition in H-atom.

3. $434.1 \text{ nm} = 4341 \text{ \AA}$ in H-atom. $\Rightarrow \frac{1}{\lambda_{\text{Photon}}} = 109677 \times 1^2 \times \left(\frac{1}{2^2} - \frac{1}{n_2^2} \right) = \frac{1}{4341 \times 10^{-10}} \Rightarrow n_2 = 5$
(Balmer series)
 $n_1 = 2$

4. $\Delta E_{3 \rightarrow 1}$ in H-atom = $13.6 \times 1^2 \left(\frac{1}{1^2} - \frac{1}{3^2} \right) \text{ eV} = 12.09 \text{ eV}$

$$\Delta P.E_{3 \rightarrow 1} = 2\Delta E_{3 \rightarrow 1} (\because P.E_n = 2E_n) = 24.18 \text{ eV}$$

5. H_β in Balmer series: $4 \rightarrow 2 : \frac{1}{\lambda} = 109677 \times 1^2 \times \left(\frac{1}{2^2} - \frac{1}{4^2} \right) \text{ cm}^{-1} \Rightarrow \lambda = 4862.76 \text{ \AA}$

$$\Rightarrow \nu = \frac{c}{\lambda} = 6.17 \times 10^{14} \text{ Hz}$$

6. Series limit for Balmer Series in H : $\infty \rightarrow 2$

$$(a) \quad \frac{1}{\lambda} = 109677 \times 1^2 \times \left(\frac{1}{2^2} - \frac{1}{\infty^2} \right) \Rightarrow \lambda = 3647.07 \text{ \AA} \quad (b) \quad I.E._H = 2.18 \times 10^{-18} \text{ J}$$

$$(c) \quad 2.18 \times 10^{-18} = \frac{hc}{\lambda} \Rightarrow \lambda = 911.8 \text{ \AA}$$

$$7. \quad I.E = 100 \text{ kcal mol}^{-1} = 100 \times \frac{10^3 \times 4.184}{6.02 \times 10^{23}} \text{ J/atom} = 6.95 \times 10^{-19} \text{ J/atom} = h\nu \Rightarrow \nu = 1.05 \times 10^{15} \text{ Hz}$$

$$8.(BC) \quad \Delta E \propto Z^2 \Rightarrow \frac{hc}{\lambda} \propto Z^2 \Rightarrow \lambda \propto \frac{1}{Z^2}$$

$$9.(C) \quad \text{As } n \uparrow, \Delta E_{b/w \text{ two shells}} \downarrow \Rightarrow \Delta E_{\max.} \text{ for } 3 \rightarrow 2$$

$$10.(C) \quad r_n \propto \frac{n^2}{Z} \Rightarrow \frac{(r_1)_H}{(r_2)_{Li^{2+}}} = \frac{Z_{Li^{2+}}}{Z_H} \Rightarrow r_{Li^{2+}} = \frac{r}{3}$$

11.(C) 4341 \AA : Visible region in H – atom (Balmer series : $n_1 = 2$)

$$\Rightarrow \frac{1}{\lambda} = RZ^2 \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \Rightarrow \frac{1}{4341 \times 10^{-8}} = 109677 \times 1^2 \times \left(\frac{1}{2^2} - \frac{1}{n_2^2} \right) \Rightarrow n_2 = 5$$

$$12.(A) \quad \frac{n(n-1)}{2} = 10 \Rightarrow n = 5; \text{ Transitions in visible region in H atom: } 5 \rightarrow 2; 4 \rightarrow 2; 3 \rightarrow 2$$

$$\Rightarrow 3 \text{ transitions}$$

$$13.(B) \quad \text{Angular Momentum (L)} = \frac{nh}{2\pi} = \frac{3h}{2\pi}$$

$$14.(B) \quad K.E_n = -E_n = 2.18 \times 10^{-18} \times \frac{Z^2}{n^2} \text{ J} = 5.45 \times 10^{-19} \text{ J} \Rightarrow n = 2$$

15.(D) $6 \rightarrow 3$: For visible region, final n should be equal to 2 in H – atom $\Rightarrow$ No transition in visible region.

16.(ABD)

17.(D) Learn it as a fact.

$$18.(BC) \quad L = \frac{nh}{2\pi}; \text{ check yourself that (B) is a correct statement.}$$

$$19.(D) \quad \Delta p \geq \frac{h}{4\pi \Delta x} \Rightarrow \text{If } \Delta x = 0; \Delta p = \infty$$

$$20.(A) \quad \lambda_{\text{particle}} = \frac{h}{\sqrt{2m \text{ K.E.}}} \Rightarrow \text{for same } \lambda : m \cdot \text{K.E.} = \text{const.} \Rightarrow \text{Lighter particle will have higher K.E.}$$

$$21.(ABC) \quad r_n \propto \frac{n^2}{Z} \quad V_n \propto \frac{Z}{n} \quad E_n \propto \frac{Z^2}{n^2} \quad L = \frac{nh}{2\pi} \quad (\text{doesn't depend on } Z)$$

$$22.(A) \quad H_\beta \text{ in lyman series} \Rightarrow h\nu = 2.18 \times 10^{-18} \times 1^2 \times \left(\frac{1}{1^2} - \frac{1}{3^2} \right) \Rightarrow \nu = 2.92 \times 10^{15} \text{ Hz}$$

$$23.(D) \quad \text{As we know } r \propto \frac{n^2}{Z}; \quad \text{For } Li^{2+}, Z=3 \Rightarrow r \text{ is minimum}$$

24.(D) As we know

$$\frac{1}{\lambda} = RZ^2 \left(\frac{1}{n_1^2} - \frac{1}{n_2^2} \right) \quad ; \quad \text{higher } Z \text{ means lower } \lambda.$$

25.(C) As we know

$$E_n \propto -13.6 \frac{Z^2}{n^2}$$

Since $E_1 = -54.4 \text{ eV}$

Hence He^+ ion.

26.(A) If 10.2 eV of energy is absorbed then electron goes from $1 \rightarrow 2$. $\Rightarrow$ Increase in angular momentum $= \frac{h}{2\pi} = 1.05 \times 10^{-34} \text{ Js}$

27.(A) Since one visible quanta is there, hence the transition must be to the second quantum number. Hence the final transition is $2 \rightarrow 1$.

28.(B) Radio wave have maximum wavelength

29.(A) Zeeman effect explains splitting in magnetic field.

EXERCISE-C

1. $\lambda_e = \frac{h}{\sqrt{2m_e K.E_e}}$

$$K.E_{e^-} = 1 \text{ eV} = 1.6 \times 10^{-19} \text{ J} \quad \Rightarrow \quad \lambda_e = \frac{6.626 \times 10^{-34}}{\sqrt{2 \times 9.1 \times 10^{-31} \times 1.6 \times 10^{-19}}} = 12.26 \text{ \AA}$$

2. $\lambda = \frac{h}{p} \quad \Rightarrow \quad \text{If } p \downarrow, \lambda \uparrow \quad \Rightarrow \quad \text{when } p_{\text{new}} = \frac{1}{2} p_{\text{old}}, \lambda_{\text{new}} = 2\lambda_{\text{old}} = 10^{-7} \text{ m} = 1000 \text{ \AA}$

3. (a) First find the no. of unpaired e^- s in each of the atom by writing their electronic configuration.

(b) Use $\mu = \sqrt{n(n+2)}$ B.M. where $n = \text{No of unpaired } e^- \text{ s.}$

4. (a) $2p_x$ or $2p_y$ (b) $4d_{z^2}$ (c) $3p_z$ (d) $3d_{xy}$ and $3d_{x^2-y^2}$ (d) $4p_x$ and $4p_y$

5. Do it yourself.

6. Use the electronic configuration table in the chapter.

7. Orbital Angular Momentum $= \sqrt{\ell(\ell+1)} \frac{h}{2\pi}$

$$\Rightarrow \quad 4s : 0; \quad 3p : \sqrt{2} \frac{h}{2\pi}; \quad 3d : \sqrt{6} \frac{h}{2\pi}$$

8. $\Delta v \geq \frac{h}{4\pi m \Delta x} = \frac{6.626 \times 10^{-34}}{4\pi \times 9.1 \times 10^{-31} \times 10^{-10}} = 5.8 \times 10^5 \text{ ms}^{-1}$

9.(B) $2s : \ell = 0 \Rightarrow \text{orbital Angular momentum } \sqrt{\ell(\ell+1)} \frac{h}{2\pi} = 0$

10.(D) $F (Z=9) : 1s^2 2s^2 2p^5$: p-orbital has $5e^-$ s.

$\text{Na} (Z=11) : 1s^2 2s^2 2p^6 3s^1$: s-orbital has $5e^-$ s

$\text{Fe}^{3+} (Z=26) : 1s^2 2s^2 2p^6 3s^2 3p^6 3d^5$: d-orbital has $5e^-$ s.

$\text{Mn} (Z=25) : 1s^2 2s^2 2p^6 3s^2 3p^6 3d^5 4s^2$: d-orbital has $5e^-$ s.

11.(C) $\lambda_{\text{Ball}} = \frac{h}{mv} = \frac{6.626 \times 10^{-34} \times 3600}{0.2 \times 5} = 2.38 \times 10^{-30} \text{ m.}$

12.(C) 3s is more closer to the nucleus. $3s > 3p > 3d$.

13.(A) Rb(Z = 37): $1s^2 2s^2 2p^6 3s^2 3p^6 4s^2 3d^{10} 4p^6 5s^1$: **Outermost e^-**

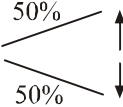
14.(A) d orbital e^- : Orbital angular momentum $= \sqrt{\ell(\ell+1)} \frac{h}{2\pi} = \sqrt{6} \hbar \left(\hbar = \frac{h}{2\pi} \right)$
 $(\ell=2)$

15.(C) Cl: $1s^2 2s^2 2p^6 3s^2 3p^5$: Unpaired e^- :

16.(A) M-shell: (n = 3) $\Rightarrow$ No. of orbitals = $n^2 = 9$

17.(B) $r = 8.46 \text{ \AA} = 0.53 \frac{n^2}{Z} \text{ \AA} \Rightarrow n = 4$

Max. e^- s in $n = 4$: $= 2n^2 = 32$

18.(C) In any shell, max. no. of e^- s = $2(2\ell + 1)$  $\Rightarrow$ Max no. of e^- s with same spin = $2\ell + 1$

19.(C) Value of $\ell < 'n'$ $-\ell \leq m \leq +\ell$, $m_s = \pm \frac{1}{2}$

20.(D) Fe^{2+} : d-orbitals have $6e^-$ s.

Na: $1s^2 2s^2 2p^6 3s^1$: 5 s-orbital e^- s

Li: $1s^2 2s^1$: 3 s-orbital e^- s

N: $1s^2 2s^2 2p^3$: 4 s-orbital e^- s

P: $1s^2 2s^2 2p^6 3s^2 3p^3$: 6 s-orbital e^- s

21.(D) Hund's rule is related to the degenerate orbitals.

22.(C) 6s will be closest.

23.(C) $2.5 \hbar$ means $\frac{5h}{3\pi}$ can not be correct value of angular momentum.

24.(D) To emit a photon, e^- has to go to lower energy level. In $n = 1$, e^- can't go to a lower energy level further.

25.(A) All half filled orbitals so are most stable.

26.(A) Zero probability will be for P_z .

27.(A) A nodal plane is the one where probability of finding electron is zero.

